





# The Oceanography Report



The Oceanography Report  
The journal of physical, chemical, geological, and biological oceanography.

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## Farewell Remarks by Chris Mooers

For 6 years it has been my privilege and pleasure to serve the Ocean Sciences Section as an elected officer. First as Secretary, then as President-Elect, and finally, as President. The fact that I have pursued through these years has been to help the Ocean Sciences Community awaken to opportunities within AGU to "flex its muscles" as a large, strong, rapidly growing segment of AGU, and to become involved in using AGU's programmatic and professional development of the ocean sciences. (Our section is the second largest of 10 sections and may become the largest within several years, if the present relative rate of growth is sustained.) Like any large organization, AGU has inertia; however, we have learned that it is flexible and yields to aid, in fact, supports the initiatives of activist. And we have only scratched the surface.

In recent years, the Ocean Sciences community has continued to grow and mature. Much of its scientific communications has involved reporting results of the first two generations of "large science" programs in AGU (and other) meetings and journals, while not neglecting advances still to be gained by "small science." Major process, especially on the mesoscale, and regional studies have yielded not only major scientific results but also the community expertise and confidence to proceed to higher levels of quantitative science, to larger-scale and longer-term problems, and to multidisciplinary studies. With declining funding levels and aging facilities, the field has become more competitive, just as it is poised to move into what may be its "prime." In this situation, some of us thought it was important to move toward a higher level of cohesiveness, professionalism, and sense of community. One avenue of initiative open to us in community building was a stronger role in the programs of AGU.

Let me summarize some of the initiatives taken. The Oceanography Section was re-named the Ocean Sciences Section to recognize the breadth of our interest and our continuing transition to a quantitative, predictive science. The *Oceanography Report*, edited by Arnold Gordon and issued on a monthly basis in *Eos* has been a great success as a vehicle for community news; it has paved the way in AGU policy circles for an Ocean Sciences Bulletin if and when we are ready for it. Under A. D. Kirwan's, and then Jim O'Brien's, editorial leadership, a component *Journal of Geophysical Research* dedicated to the oceans has been brought into existence. The recently published *Careers in Oceanography* booklet, produced under Charles Hollister's leadership, has given us an exciting, straight-talking recruiting pamphlet for the first time. We now have an Ocean Sciences Education Secretary, Peter Brewer, to answer inquiries from young people. Ocean Sciences luncheons have become fixtures at all national meetings. They are used for building a sense of community through discussion of Section and AGU issues; for presentation of a major, informal talk by a community leader, usually from Washington, as an unofficial exercise in accountability; and for presentation of the recently instituted Ocean Science Awards, given for outstanding and sustained contributions to the community through service, research, leadership, and so forth. A campaign has continued to welcome biological oceanographers at AGU meetings and as AGU members. Part and parcel with this has been the continued series of joint, experimental meet-

ings with ASLO, and now other societies, beginning with San Antonio (convened by Worth Nowlin and Dick Eppley) in February 1982, collocation of ASLO and AGU meetings in December 1982, the Ocean Sciences Meeting 1984 in New Orleans in January (convened by John Apel and Dick Barber), and the upcoming integrated ASLO/Ocean Sciences Section Program (planned by Wolfgang and Pat Kremer) at the December 1984 AGU Meeting. We have now settled on a biennial Ocean Sciences Meeting.

In recent years, our hardworking, creative program chairmen for AGU national meetings have been Clayton Paulson, John Bane, Dave Cutchin, and Bob Molnar. Many others have served on program committees, as session chairmen, on ad hoc committees of the Section, and on the standing committees of the AGU per se. The point to be emphasized is that the Union and the Section are very much participatory entities and that the Ocean Sciences community can be strengthened by even broader participation. Please let the new Ocean Sciences President, Joe Reid, know if you want to be involved and in what capacity.

Part of the AGU-wide effort to update and "standardize" section bylaws is the establishment of an Ocean Sciences Executive Committee, consisting of three elected section officers and up to five others. Under my "reign," besides Joe Reid, Peter Brewer, and me, the members have been Arnold Gordon, Jim O'Brien, John Apel, Harmon Craig, and Jim Baker. This Executive Committee does strategic planning for the Section and effectively extends the leadership. Together with the Geodesy Section and the President of AGU, we have pressed for the issue of a national policy statement on the releasability of GEOSAT data. On a Union-wide basis, the role of the sections in the selection of AGU Fellows has been greatly strengthened. The Ocean Sciences Section now has an ad hoc AGU Fellows Nominating Committee, chaired by the President-Elect. Consequently, we have presented better documented cases and have been much more successful in the election of Ocean Sciences Section members as AGU Fellows. (There is still a need for more members to nominate colleagues.) Last, a new AGU monograph series in coastal and estuarine regimes has been initiated, with the first volumes to appear within a year.

Much more potential remains to be tapped within AGU. For example, the Ocean Sciences Section has not exploited much the topical meetings mechanism for scientific communications. We have not been fully active in all-Union sessions (e.g., Frontiers of Geophysics). With more participation in organizing scientific meetings, the quality and coherence of our talk and poster sessions can be improved. (We have made improvements there through the aggressive pursuit of better meeting rooms and the organization of pre-designated thematic sessions.) More could be done with joint sessions with other sections; such as, Geodesy, Hydrology, and Atmospheric Sciences. AGU can help us in organizing more effective graduate student recruitment efforts. The Ocean Sciences Bulletin idea, modeled after the *Bulletin of the AMS* and *Physics Today*, is waiting for someone to come forward as the inaugural editor. There is also room for more specialty journals. Finally, the possibility exists to organize an Ocean Sciences society within AGU if there is sufficient need, interest, consensus, and leadership.

There are exciting times ahead for the ocean sciences, and the Ocean Sciences Section needs to anticipate them, and to help lead them. For example, owing to the imperatives of the scientific agenda for global ocean circulation, marine ecosystem, biogeochemical transport, benthic province, and other studies, of technological opportunities provided by super computers, microprocessors, new sensor systems, future ocean satellites, retrievable and expendable profilers, moored and drifting buoys, and ships-of-opportunity, of scientific opportunities provided by new understanding and models, and of programmatic needs associated with the missions of the new National Ocean Service and the soon-to-be revitalized Naval Oceanography Program, the dawning of global synoptic oceanography is at hand. The community will need help in adjusting to the sociological shock of working in real-time, and the Ocean Sciences Section can help to communicate and foster the "revolution," and the concomitant educational needs and employment opportunities that the new industry of operational oceanography will bring. If oceanographers can commit themselves to work in real-time, benefits from the operational oceanography can accrue in the form of cooperation and resources to achieve otherwise unobtainable goals in global ocean studies of climate variability and related topics. Other benefits include the ability to conduct up-to-date quality control of data acquisition, adaptive sampling strategies, and onboard scientific analyses. Similarly, the prospective large-scale exploration for resource management (exploitation and conservation) of our EEZ will provide new scientific challenges and op-

portunities in the coastal ocean, as well working with other nations in developing their ocean science capabilities for EEZ resource management. For all of this, oceanographers will need to communicate with other science disciplines, engineers, and technologists. Again, the Ocean Sciences Section (or Society) can serve to help traverse these new frontiers.

Chris Mooers is outgoing President, AGU Ocean Sciences Section.

## News & Announcements

### Physics of Shallow Estuaries and Bays

The objective of the symposium is to promote the exchange of information on recent developments in estuarine physics between physical oceanographers and coastal engineers. Emphasis will be on shallow estuaries and bays (well-mixed and partially mixed) and reef lagoons.

The symposium stresses results of field and laboratory measurements and the formulation of the governing equations and boundary conditions as opposed to numerical techniques. Both well-mixed and partially mixed estuaries are included. The following topics will be covered: (1) large-scale transport processes (mixing); (2) tide and wind-induced residual flow; (3) small-scale turbulence; (4) suspended sediment transport; and (5) exchange between ocean and estuary/bay.

The symposium will be held at the Rosenstiel School of Marine and Atmospheric Science, University of Miami, Florida. The dates of the symposium, August 29, 30, and 31, 1984, were selected to allow participants to combine attendance of the symposium with the 19th International Conference on Coastal Engineering (ICCE) to be held in Houston, Texas, September 3-7, 1984. The symposium is cosponsored by the Coastal Engineering Research Council of the American Society of Civil Engineers and the Rosenstiel School of Marine and Atmospheric Science.

For additional information, registration, etc., write: Physics of Shallow Estuaries and Bays, c/o Division of Ocean Engineering and Applied Marine Science, RSMAS, University of Miami, 4600 Rickenbacker Causeway, Miami, FL 33149 (telephone 305-361-4161).

## Meetings

### New Directions in Internal Wave and Microstructure Research

Collaboration of oceanographers across traditional boundaries always results in exciting new insights into the complex processes which govern oceanic mixing. Such collaboration was evident at the second annual 'Alu Hui' Hawaiian Winter Workshop, held in Honolulu, Hawaii, January 18-20, 1984. This year's topic was Internal Gravity Waves and Small-Scale Turbulence. Participants from Europe, Canada, and the United States reviewed recent developments and proposed intriguing studies in the kinematics and dynamics of internal waves, fine structure, and microstructure. Here we summarize the highlights of the meeting and identify some of the emerging trends, all subject to the participants' biases.

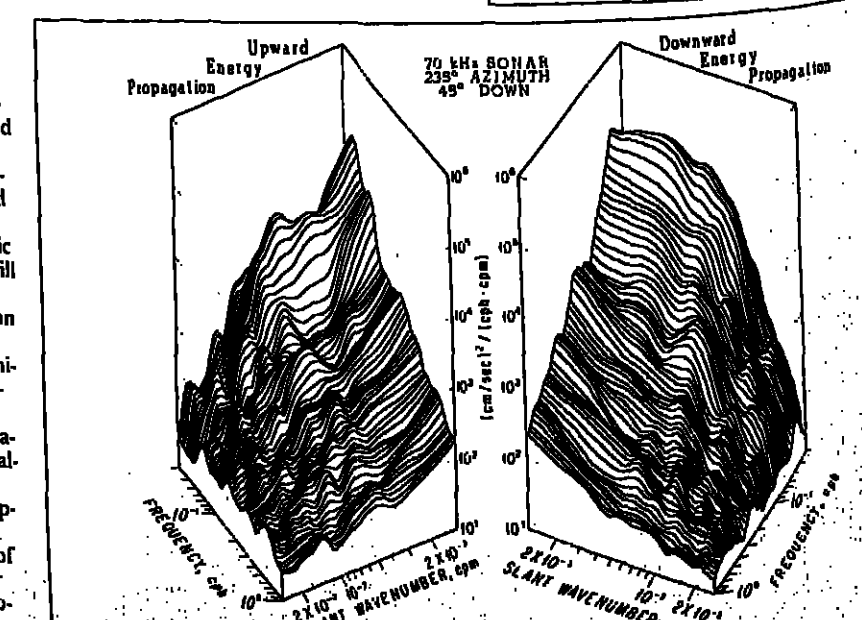


Fig. 2. Wave number-frequency spectrum of the upward and downward propagating internal wave energy. The spectra were measured with a Doppler sonar off the coast of southern California. The spectrum is representative of the internal wave field in the depth range from 80 to 600 m and over a period of 18 days (courtesy of R. Pinkel).

## Internal Waves

Internal waves are viewed as an important link in the overall oceanic energy cascade from the large scales of generation to the small scales of dissipation. Although the dominant sources and sinks for internal waves have not been identified, the following concept is generally accepted: Energy enters the internal wave field at large scales and cascades down to small scales by nonlinear wave interactions. When the shear reaches a critical value, the waves break and generate small-scale turbulence and microstructure. At microscale energy is dissipated by molecular processes. Research that led to this picture has been dominated by the concept of a universal internal wave spectrum, an idea introduced over a decade ago by *Ginart and Munk* (1972). During the workshop, the concept of a universal spectrum was challenged, whereas the link between internal waves and microstructure was substantiated.

"Universal" Spectrum. Observed spectra usually fit the "universal" spectrum to within a factor of 3 for frequencies significantly above the internal and less so in the near-inertial band (*Wunsch, 1976; Briscoe; Levine*). (Note: Undated references refer to talks given at the workshop. These talks will be published in the proceedings. Copies may be obtained from Peter Müller, University of Hawaii, Department of Oceanography, 1000 Pope Road, Honolulu, HI 96822.) This means that there is an order of magnitude variation in the spectral levels. These variations are likely the dynamical signatures of the sources, sinks, and internal transfers of the internal wave field. It is these dynamical features that have become the object of internal wave research.

The deviations of the internal wave field from the universal form exhibit definite patterns. Energy in the near-inertial frequency band varies in response to storms and to mesoscale features (*D'Asaro*), as well as geographically and with depth (*Fin, 1981*). The energy in the higher frequency continuum varies seasonally and geographically (*Briscoe, Levine*, and *Figure 1*) and near topographic features (*Eriksson*). Clear patterns exist in data; explanation of their dynamics is a challenge for future research.

New measurement techniques like Doppler sonars reveal local spectra that are not smooth but show an irregular structure with ridges and shoulders (*Pinkel and Figure 2*). Nonlinear interactions among internal waves have been analyzed by using the weak resonant interaction approach. Detailed calculations had

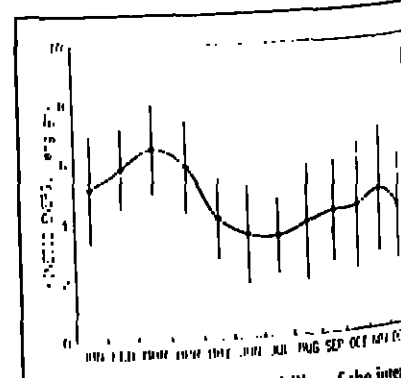


Fig. 1. Seasonal variability of the internal wave continuum. The data points represent the high frequency (0.1-2.0 cph) horizontal kinetic energy from 13 moorings in the North Atlantic over a 10-year period. The moorings are all subsurface. The instrument depths are 114-508 m, and the locations range from 18°-30°N and 12°-71°W. The error bars are plus and minus one standard deviation (courtesy of M. Briscoe).

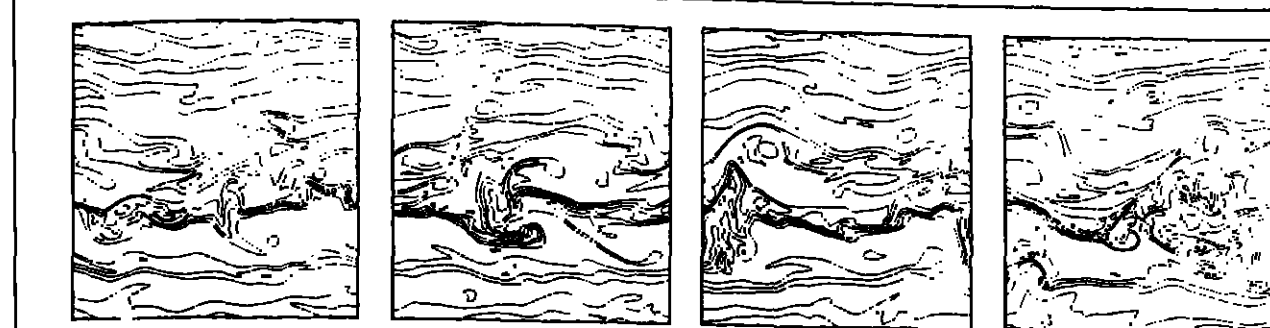


Fig. 3. Overturning of density surfaces in a two-dimensional simulation of strongly interacting internal waves. The domain represents a vertical plane, the lines isopycnals. The Richardson number of the flow is about 0.7. The frames are separated by about half a buoyancy period (courtesy of G. Holloway).

been made by using this technique, and an inertial range theory has emerged similar to the one in turbulence theory (*Alcorno and Müller, 1981*). Nonlinear interactions cascade energy down the spectrum from the generation to the dissipation scale. The level of the energy spectrum adjusts itself to the energy flux through the spectrum. The downscale cascade is associated with an energy transfer from high to low frequencies, somewhat opposite to conventional wisdom. The appropriateness of the weak resonant interaction approach for small-scale waves has been questioned because interaction times are often much shorter than the periods of the waves (*Holloway, 1980*).

Now, nonlinear interactions are investigated by two new methods: numerical integration of the Navier-Stokes equations in two dimensions (*Holloway and Figure 3*) and Monte-Carlo simulation of the Eikonal equations

that describe the evolution of a small-scale wave in a background wave field (*Heney*). Their preliminary calculations show unexpected and exciting results: an upward mass flux (mixing) at low wave numbers in *Holloway's* calculation and preferred layers of breaking "patches" in the Eikonal approach. Unlike the weak interaction calculations, the new approaches produce space-time results that will eventually allow direct comparison with experimental data. Such comparison certainly will stimulate a greater interaction between theoretical and observational oceanographers.

## Sources and Sinks

Numerous sources and sinks have been proposed for the internal wave field (see, e.g., *Olters, 1983*). Observationally, the situation might be summarized as "a little bit of evidence for everything" (*Briscoe*). No dominant generation or dissipation mechanism has been identified, although some progress is occurring on some mechanisms.

Theoretical and observational evidence is emerging that the wind generates near-inertial frequency waves at large vertical scales (*D'Asaro*) and that internal waves and the mesoscale flow strongly interact (*Watson*). Classically, it has been assumed that the internal wave field dissipates its energy predominantly in the interior of the ocean, through small-scale turbulence. Calculations (*Eriksson*) indicate, however, that the loss of internal wave energy at a sloping boundary might be substantial and could be the major energy sink of internal waves. Significant sinks of energy also may occur in critical layers when near-inertial waves become trapped within fronts or eddies (*Kunze and Sanford, 1984*). These losses would be concentrated at particular locations in the ocean and not spread uniformly throughout its volume.

## Current Fine Structure

Existing velocity and temperature measurements clearly show that linear internal waves alone cannot explain all of the observed structure within the internal wave frequency band (*Müller et al., 1978*). In particular, the coherence between current meters as a function of vertical separation drops rapidly within the first few meters and then decays more slowly on a scale of many tens of meters. The rapid drop is traditionally attributed to current fine structure. At frequencies well above inertial frequency, current fine structure has an energy density comparable to that of internal wave motions. The kinematical and dynamical character of current fine structure is unclear. The traditional view is that it represents internal wave currents concentrated at small vertical scales because of the fine structure in the Brunt-Väisälä profile. A different view (*Müller*) holds that current fine structure is an entirely different type of motion with well-defined, distinct dynamics. Unlike internal waves, this "vortical mode" of motion (*Riley et al., 1981*) carries potential vorticity. Current fine structure might, hence, be the small-scale realization of the same mode that represents quasigeostrophic flows at mesoscales.

The separation of internal waves and "vortical" motions is also a problem in the atmosphere. In meteorology, the "vortical" mode is called "stratified two-dimensional (2-D) turbulence." The observed atmospheric mesoscale spectra are roughly consistent with theories of upscale integral ranges in stratified 2-D turbulence (*Lilly*).

The implications of the existence of the vortical mode for the dynamics have not yet been explored, but we expect the vortical mode to be intimately connected and intertwined with the internal gravity mode of motion (*Holloway*). A distinction between the internal gravity and vortical mode of motion requires the measurement of vorticity on small scales, a measurement that to date has not been possible because of lack of suitable instruments; however, a "vorticity meter" is now being developed by Sanford (personal communication, 1984), so that such distinction might soon become possible.

## Small-Scale Turbulence

If double diffusive effects are ignored, small-scale turbulence measurements are almost always discussed within the following,

now classical, framework: Estimates of oceanic mixing rates can be made from velocity and temperature measurements that resolve the small scales on which molecular dissipation occurs. Measurements of the Richardson number on the meter scale commonly show values of the order of 1 or less, suggesting that shear instability is a major mechanism for mixing. Assuming that the shear responsible for the small values of the Richardson number is due to internal waves, dissipation driven by shear instability becomes an energy sink for the internal wave field. Accordingly, the rate of mixing and the properties of the internal wave field are related. If this link were understood, the rate of mixing could be parameterized in terms of the energy sources and environmental parameters of the internal wave field.

Studies of the relationship between small-scale turbulence and the internal wave field clearly require measurements of both the turbulence, using microstructure instruments, and the internal wave shear and density fields, using larger scale measurements. Existing evidence suggests that the internal wave field is highly random, so many measurements are required. Instrument systems capable of repeated measurements of both microstructure and internal wave scales have only recently become available (*Gregg, Osborn*) and are limited to use in the upper few hundred meters. Simultaneous measurements of the internal wave spectrum and microstructure over periods long enough for significant changes in the internal wave field to occur are not yet available. It is, therefore, not surprising that the research in this field is still exploratory.

If, as hypothesized, small-scale turbulence is driven by the internal wave field, its structure should reflect the structure of the wave field. Observationally, this issue is complicated by the possibility of turbulence caused by double diffusion which is ignored here. Nevertheless, several promising links between the internal wave field and oceanic turbulence are emerging.

*Patchiness.* Measurements of small-scale turbulence generally show that the individual mixing events are not randomly distributed, but concentrated into "patches" of high activity. These "patches" vary in size from centimeters to 10-20 m (*Gregg, Dugan, Osborn, and Figure 4*), with the smaller patches being more common.

Two theoretical approaches based on internal waves predict such a structure: Calculations of the vertical distribution of the Richardson number, *Ri*, made assuming a Garrett and Munk internal wave spectrum, and Gaussian statistics (*Desaubies*). If a turbulent patch is assumed to occur whenever *Ri* < 1/4, a range of "patch" sizes, comparable to that observed, is computed. A more detailed comparison with the observed "patch" statistics would be interesting. Eikonal calculations,

which trace individual waves in a background internal wave spectrum, have also been used to model the spatial distribution of small-scale turbulence (*Heney*). It is assumed that an individual wave breaks when it reaches a sufficiently high wave number. One such calculation shows the persistent clustering of the breaking events at a particular level, suggesting the formation of a turbulent "patch." This calculation suggests that it may be possible to formulate general criteria for the location of such patches as a function of the background shear field.

Recent experimental work suggests a link between the larger patches and near-inertial frequency shear. One such patch, which persisted for nearly a day, occurred at the same depth as a small inertial jet (*Gregg*).

*Shear Statistics.* The universal internal wave spectra have dominantly been energy spectra and have not accurately described the statistics of the shear and density gradient fields, particularly on scales smaller than 10 m. Such a description is needed if accurate models of the link between microstructure and internal waves are to be developed. A "universal" shear spectrum has been proposed by *Gregg et al. (1981)*, but it is not complete. There is still uncertainty as to whether the shear at 10 m scales is dominantly inertial, as at larger scales (*D'Asaro*) or dominantly high frequency (*Pinkel*). This particular question is complicated by Doppler-shifting of small-scale velocity features. Basic descriptive work is needed on the shear and density gradient (distribution, spectrally, spatially, and temporally).

*Kelvin-Helmholtz Billows.* A variety of ingenious arguments developed in the last decade allow *K<sub>h</sub>*, the vertical diffusivity for mass, to be estimated from microstructure parameters; however, a clear picture of the three-dimensional structure and evolution of these mixing events has not yet emerged. Generally, mixing is envisioned as being caused by Kelvin-Helmholtz billows. The structure of these billows was very nicely depicted by the artist of the Roman scallop shell mosaic shown on the cover. These billows have been extensively studied in the laboratory and have been observed at one location in the upper ocean (*Hooks, 1968*). The extrapolation of laboratory studies to the ocean may be difficult, due to side-wall effects in the laboratory studies (*Thorpe*). A variety of other stratified shear flow instabilities with structures distinct from Kelvin-Helmholtz billows, such as wave breaking and critical layer absorption, have been observed in the laboratory (*Thorpe, 1973*) and may also occur in the ocean. Turbulence research in other fields has benefited greatly from flow visualization studies that aim to identify the dominant structures of the turbulent flow. Once the structures of a flow have been identified in this way they can usually be identified in point measurements. It seems likely that similar studies using dye or high-frequency neocuties would likewise increase our understanding of oceanic turbulence.

On the theoretical side, the nonlinear stability of stratified flows has been investigated by using a constrained energy method due to V. I. Arnold (*Abbaranel*). Applied to the customary parallel shear flow in the presence of stratification, one proves that the flow is nonlinearly stable for Richardson numbers greater than unity. The theoretically interesting regime is hence  $1/4 < Ri < 1$  where the flow is stable to infinitesimal perturbations, but may be unstable to finite perturbations. Shear-generated turbulence in homogeneous fluids is being studied with new second-order closures. Employed in numerical models, these yield good predictions of the observed Reynolds stress tensors evolving in strained and sheared wind tunnel flows (*Gallagher*).

Oceanography (cont. on p. 380)

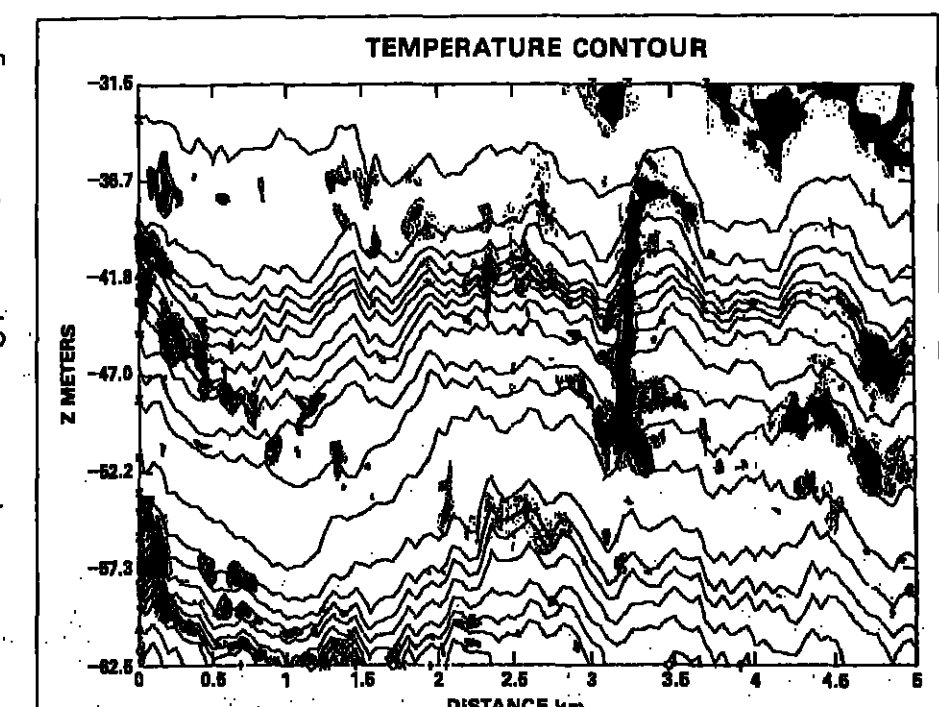


Fig. 4. Patches of microstructure in a temperature section. In the shaded areas the microstructure activity is 30% above the mean. The lines represent isotherms. The data are from the Sargasso Sea and were obtained with a thermistor chain towed at a speed of about 250 cm s<sup>-1</sup> (courtesy of J. Dugan).



## Oceanography (cont. from p. 379)

## Parameterization

Many oceanographers prefer to study the large-scale motions of the ocean. These oceanographers regard internal waves and small-scale turbulence as subgrid-scale noise and ask for the parameterization of subgrid fluxes in terms of large-scale flow characteristics. They ask for eddy diffusion and viscosity coefficients. Here the state of affairs is still not satisfactory. Most work on parameterization has been concerned with the vertical diffusion coefficient  $K_v$ . A typical value of  $0.1 \text{ cm}^2 \text{ s}^{-1}$  seems not to be inconsistent with microstructure measurements and the kinematics and dynamics of internal waves (Garrett). A similar value is obtained when the observed large-scale hydrographic field is fitted by beta-spiral methods (Olbers) but that value includes an artificial contribution due to averaging of the data. Basic questions are still unanswered. For example, how much of the vertical mixing is done in the interior of the ocean and how much is done in boundary layers? Does the value  $K_v$  have a strong depth dependence? The answers to these questions could have dramatic implications. Changing the depth dependence of the dissipation rate changes the direction of the meridional circulation in an advective diffusive model of the thermohaline circulation (Gargett). The momentum fluxes are even less established. There are only spotty measurements.

Some of them imply significant eddy viscosity coefficients [e.g., Brown and Owens, 1981], but no coherent picture has emerged from the sparse data.

## Conclusions and Trends

Internal gravity waves and small-scale turbulence are the motions by which the ocean mixes momentum and mass. The specific way in which this mixing is done has pronounced effects on geostrophic eddies and the general circulation. To understand these grander scales of motions, we must understand the smaller-scale mixing processes. Internal wave research is presently undergoing a transition from a dominantly kinematic study of spectral slopes to a dominantly dynamic study of sources, sinks, and internal fluxes. The link between internal waves and oceanic turbulence is becoming more apparent, and the glimmers of a dynamical understanding are emerging. The parameterization of the internal wave and turbulent fluxes, which is a major goal of these studies, has not yet been achieved. Further progress will come from simultaneous measurements of internal waves and microstructure and from a detailed comparison of experimental data with the results of numerical models. These experiments and studies require collaboration of oceanographers across specific areas of interest, a beginning of which was witnessed at the Hawaiian Winter Workshop.

## Acknowledgments

The second 'Aha Huli' Hawaiian Winter Workshop on Internal Gravity Waves and Small-Scale Turbulence was supported by the Office of Naval Research, Hawaiian Institute of Geophysics contribution 1483. We thank all the participants of the workshop for their valuable ideas (many of which have gone unacknowledged here), and for their permission to quote unpublished materials.

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designed for applications to various hydrological modeling and data analysis problems, including precipitation; (3) that 1-3 day synoptic, workshops, or special sessions at AGU/AMS/ASA or other conferences are organized; (4) that AGU publish monographs containing review papers; (5) that special issues of appropriate journals (addressing important recent developments in precipitation) are published; and (6) that research projects having strong interdisciplinary makeup are funded by NOAA/NSF/NASA/USGS and other agencies.

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We therefore recommend that the hydrological community actively participate in the planning of the experiment and in the analysis of the experimental data for hydrological studies. For this purpose, immediate actions are needed, first, to coordinate the planning of a network for collecting hydrologic data on basins in the central United States (e.g., streamflows, groundwater levels, soil moisture, evaporation, and evapotranspiration) and, second, to provide input to the STORM Central planning team for coordinating the collection of these data in conjunction with meteorologic data.

In addition to the above, we would like to make the following recommendations: (1) that the Office of Hydrology of the NWS establish a working group to study the users' need for quantitative precipitation forecasts; (2) that the National Climatic Data Center establish a hydrometeorological data archive.

**Plan on San Francisco Now!**  
**AGU Fall Meeting**  
**December 3-7, 1984**

## News

## Thunder Day Increase

A report issued by the Illinois State Water Survey concludes that annual values of thunder days for North America exhibited a general increase of about 15% from 1901 to 1945, followed by a general decrease of 10% from 1945 to 1980. A study of the variability of thunder days across North America showed a general decrease with time, particularly after 1940. A major finding of this study is that frequencies of thunderstorms over areas as large as the North American continent show major long-term trends.

The report, "Temporal Distribution of Global Thunder Days," summarizes the results of a 1-year study by Stanley A. Changnon, Jr., and Chin-Fai Hsu of the temporal variations of thunder-day records during 1901-1980 using quality weather records from weather stations scattered around the globe. A thunder day is recorded when one or more peals of thunder are heard anytime during the 24-hour period from midnight to midnight, which is consistent with the definition of a thunderstorm used at first-order weather stations since 1897. They found most stations in the northern hemisphere north of 45° latitude exhibited a general increase in thunder activity from 1901 to 1980. The project was funded by the National Science Foundation.

Data for the study was collected from 90 stations in North America and 131 stations elsewhere around the world. The stations in North America selected were those that had operated on a 24-hour basis for the period of 1901-1980. The data from most foreign stations had periods of record of only 30-50 years, which limited the investigations of trends in thunder days in areas outside of North America. The data were evaluated to check for shifts in thunder-day frequencies due to station relocation, for major discontinuities between station averages and area averages, and for possible changes in thunder frequencies due to potential noise problems, such as might be caused by increased air traffic near a station. These tests led to the conclusion that four stations (Atlanta, Ga.; Kansas City, Mo.; Portland, Maine; and Toronto, Ont., Canada) had records that were incorrect and so were not used. These tests were not applied to the foreign stations because of a lack of station histories.

Examination of the temporal distributions of thunderstorms was done in several ways, including factor analysis and other statistical tests. The continent was first divided into four sectors with divisions at 100° W longitude and 40° N latitude, and temporal distributions in each sector were determined. Analysis confirmed that there were major regional differences in temporal frequencies of thunder days across North America. General increases in frequencies from 1901 to 1980 were found in the northwest and northeastern sectors. The southwest sector frequencies generally increased in the first 45 years and then rapidly decreased in the following 10 years. Relatively low frequencies were experienced from 1956 to 1980. The southeast sector exhibited little trend until 1950, followed by a general decrease until 1980.

Stations with similar 5-year average values were classified into regions of similarity. This classification method produced 14 regions in the United States and Canada, which themselves fell into two major regions. One was the northern United States and all of Canada, where all stations exhibited a general increase in thunder-day frequency, and the other was the southeastern third of the United States, which exhibited a major decrease in thunder-day frequencies after 1925. A narrow transition zone separated the two major regions.

The 14 regions in North America can be related to major climatic zones such as differ-

ences in major air mass sources, particularly in summer, areas of cyclogenesis, and location of major cyclone tracks, the southeastern United States, for which there was a major decrease in thunder-day frequency, is an area of major industrial development and pollution in the United States, and the decrease may or may not be related to possible man-made influences.

The report concludes that the major temporal variations in thunder days were largely due to major shifts in the atmospheric circulation, reflected in the continental scale frequencies of cyclonic storms. Historical regional data on cyclone frequencies in North America would be useful in identifying the regional trends found. Comparison of thunder days with actual thunderstorm durations and other measures of atmospheric electricity are needed to help assess any temporal fluctuations in lightning discharge activity.

This news item was contributed by Steven D. Hilberg, Extension Services Coordinator, Illinois State Water Survey, Champaign, IL 61820.

## MORP: Keeping Track of Meteorites

If one considers the level of significance frequently granted to observations of small portions of an individual meteorite sample, the question of sampling error arises. A project to answer this question has been underway in Canada for 9 years (1974-1983). The Meteorite Observation and Recovery Project, or more simply MORP, is an observational network designed to evaluate the frequency of meteorite falls on the surface of the earth [see I. Halliday, A. T. Blackwell, and A. A. Griffin, *J. R. Astron. Soc. Can.*, **72**, 15, 1978]. Other major networks for observation have been operative in the United States and Central Europe. Halliday, Blackwell, and Griffin reported recently on MORP results: "...the total mass deposited on the ground is 142 kg year<sup>-1</sup> in 10<sup>6</sup> km<sup>2</sup>. ...it is obvious that a very small portion of the potential harvest is ever located" [*Science*, **223**, 1405-1407, 1984]. The actual numbers from the analysis are staggering. The normal recovery rate of meteorites per year in the world is no more than one or two dozen (excluding "fossil" meteorites recovered from the Antarctic glacial ice). The number of meteorite falls over the entire earth's surface is about 25,000 per year, according to the MORP analysis.

The basis of the analysis is photographic observation of fireballs in an area of clear sky. Meteorites were observed by at least two camera stations at an altitude of at least 8°. The MORP results turned out to be an observation of 1.16 x 10<sup>16</sup> kg<sup>2</sup> hours over the 9 year period. Analysis of the photographs were used to calculate masses and categorize the meteorites. Halliday et al. used as calibration, the observation and recovery of the Innisfree fall [*Meteorites*, **16**, 153, 1981].

Figures predicted by meteor sample mass give a clear picture of the sampling problem. In North America, the annual numbers of meteorite falls for total masses of 0.1, 1, and 10 kg per fall was 920, 980, and 38, for the entire land area of the earth the annual numbers for the three mass categories were 5800, 1200 and 240. Other calculations could be made to estimate the number of ore objects that survive the earth's atmosphere and reach the surface. Further, the estimates could be extended to encompass the earth-moon system of the solar system.—PMB

## Research Roundtable

A blue-ribbon panel of government, university, and industry leaders has been established to explore alternative approaches to strengthening their relationships. The 18 member Government-University-Industry-Research Roundtable Council, organized as an independent unit under the aegis of the council of the National Academy of Sciences (NAS), will address issues affecting and limiting the vitality of science in the United States. Thirteen of the 18 members are from universities and industry; 5 are senior federal officials. Dale R. Corson, president emeritus of Cornell University, is chairman of the roundtable council.

The council is the overall guiding group for the ongoing research roundtable, according to Don I. Phillips, council executive director. The council, which plans to meet three times a year, will spawn ongoing activities (including workshops, working groups, and special studies) to keep communication open among relevant groups.

At its first meeting on May 17 and 18, the roundtable council agreed on a broad and comprehensive framework of issues. The theme overriding the framework is the need to establish, strengthen, and maintain a network among relevant groups to keep the lines of communication open. The research round-

table will focus initially on three issues: the relationship of science and technology to economic competitiveness; the renewal of scientific and technical institutions (including upgrading facilities and equipment, enhancing technology transfer, and the new partnerships between universities and state governments and between universities and industry); and the enhancement of communications between scientists, industry employees, and government employees at the working level.

The roundtable was formed following the identification by several NAS panels of the need for an ongoing, neutral body to improve the communication and mutual understanding among those who fund research, those who carry it out, and those who use the results. NAS provided initial funding. Additional support has been made available by the Alfred P. Sloan Foundation and the Andrew W. Mellon Foundation.

In addition to chairman Corson, other members of the roundtable council are William G. Anlyan, Chancellor for Health Affairs, Duke University Medical Center; Kenneth J. Arrow, Joan Kenney Professor of Economics, Stanford University; Marvin Cohen, Professor of Physics, University of California, Berkeley; Edward G. Jefferson, Chairman of the Board, E.I. duPont de Nemours & Co., Inc.; Sol Linowitz, Covident Brothers Law Firm, Washington, D.C.; George E. Fike, Vice President for Corporate Research, Xerox Research Center; Alexander Rich, Sedgwick Professor of Biophysics, Massachusetts Institute of Technology; Howard A. Schneiderman, Senior Vice President for Research and Development, Monsanto Co.; Harold Shapiro, President, University of Michigan; Robert L. Sproull, President, University of Rochester; and Linda S. Wilson, Associate Vice Chancellor for Research, University of Illinois, Urbana.

Participating federal officials are Richard D. DeLauer, Undersecretary of Defense for Research and Engineering, U.S. Department of Defense; George A. Keyworth, II, Director of the Office of Science and Technology Policy; Edward A. Knapp, Director, National Science Foundation; Alvin W. Trivelpiece, Director, Office of Energy Research, U.S. Department of Energy; and James B. Wyngaarden, Director, National Institutes of Health. Frank Press, NAS President, is an ex officio member.

The next meeting of the research roundtable council is scheduled for November 29 and 30.

## Viking Lander 1: New Exhibit

The National Aeronautics and Space Administration (NASA) transferred ownership of the Viking Lander 1, which landed on Mars in 1976, to the Smithsonian Institution's National Air and Space Museum.

Requested by museum director Walter Boyne, the transfer includes the loan of the official Viking Lander plaque, which names the lander the Thomas A. Mutch Memorial Station in memory of the Viking Lander imaging team leader and NASA associate administrator for space science. Mutch died in a climbing accident in the Himalayas in 1980. The plaque is scheduled to be placed on Mars by U.S. astronauts at some indefinite time. NASA retains reclamation rights of the lander for scientific purposes.

This is the first time that a museum will own an object located on another planet. Tourist visitation at the new exhibit is not yet expected to rival traffic at the parent museum.

## U.K. Radio Science Reviews Available

Coinciding with its triennial general assemblies, the International Union of Radio Science (URSI) publishes an international review of the most significant scientific developments over the previous 3 years in the nine subject areas covered by URSI's commissions. To produce this review, international editors distill reviews from each member country of national scientific developments. For those scientists who wish to know more details about the significant scientific developments in radio science in the United Kingdom from 1981 to 1984, the British National Committee for Radio Science has made its reviews available.

Unless otherwise noted, the following surveys are available from the Royal Society, 6 Carlton House Terrace, London SW1Y 5AG, Attention: C.R. Argent.

Commission A: Electromagnetic meteorology, by J. M. Steele; available from J. M. Steele, National Physical Laboratory, Teddington, Middlesex, TW11 0LW.

Commission B: Fields and waves, by P. J. B. Claret.

Commission E: Interference environment, by F. Horner.

Commission F: Remote sensing and wave propagation, by E. D. R. Shearman.

Commission G: Ionospheric radio and propagation, by P. A. Smith; available from J. W. King, Rutherford Appleton Laboratory, Chilton, Didcot, Oxfordshire OX11 0QX.

Commission H: Waves and plasmas, by P. A. Smith; available from J. W. King, Rutherford Appleton Laboratory, Chilton, Didcot, Oxfordshire OX11 0QX.

Commission J: Radio astronomy, by D. H. Martin; available from D. H. Martin, Department of Physics, Queen Mary College, Mile End Road, London E1 4NS.

## Geophysicists

Recently elected as officers of the Brazilian Hydrology and Water Resources Association are Antonio Carlos Taffel Hoffa, president; Antonio Eduardo Leão Lanna, vice president; and Ciro Loureiro Rocha, Gilberto Valente Canali, and Orlando Vignoli Filho, directors. The new mailing and secretariat address is Associação Brasileira de Hidrologia e Recursos Hídricos—ABRH, A/C Gilberto Valente Canali, ELETROSUL/DIVH, Rua Deputado Antonio Edu Vieira, s/no—Pantanal, 88000 Florianópolis, SC, Brasil.

## In Memoriam

Johannes Theodor Thijssen, 91, died April 30, 1984. An AGU Life Fellow and a member of the Hydrology Section, he joined AGU in 1948.

## Recent Ph.D.'s

Eos periodically lists information on recently accepted doctoral dissertations in the disciplines of geophysics. Faculty members are invited to submit the following information on institution letterhead, above the signature of the faculty advisor or department chairman:

- (1) the dissertation title,
- (2) author's name,
- (3) name of the degree-granting department and institution,
- (4) faculty advisor.

(5) month and year degree was awarded.

If possible, include the current address and telephone number of the degree recipient (this information will not be published). Dissertations with order numbers, and many of the others listed, are available from University Microfilms International, Dissertation Copies, P.O. Box 1764, Ann Arbor, MI 48106.

## Iodide Enhanced Electron-Capture Detection of Halocarbons with Application to Atmospheric Methyl Chloride, Gwendolyn Louise Ball, Univ. of Mich., 1983 (GAX84-02238).

The Kinetics of Sorption Reactions at the Geothetic-Aqueous Interface (Disorption, Rate), Kerry Kipphart, Univ. of Notre Dame, 1983, (GAX84-03054).

Kinetic Studies of the Thermal Decomposition of Methylpropanoate and of Ozonolysis Reactions, Abraham Baluta, Penn. State Univ., 1983 (GAX83-27470).

Mariann Paluchiewicz, Susan Elaine Postawko, Univ. of Mich., 1983 (GAX84-02560).

Area- and Time-Dependent Motions in the Blake Escarpment Region, David Yueckehung Lal, Univ. of Rhode Island, 1983 (GAX84-01410).

Measurements and Analysis of Diffuse Solar Irradiance, Clifford Bruce Baker, Univ. of Mich., 1983 (GAX84-02235).

Micene Stable Isotope Stratigraphy and Paleogeography (New Zealand, Spain, Pacific Ocean), Tom Stuart Louie, Univ. of Rhode Island, 1981 (GAX83-28479).

A Multi-Tracer Study of the Abyssal Water Columns of the Deep Bering Sea, Including Sediment Interactions, II, A Six Zone Regionalized Model for Bomb Radiocarbon and Carbon Dioxide in the Upper Kilometer of the Pacific Ocean, John R. Toggweiler, Columbia Univ., 1983 (GAX83-27510).

Neotectonics of the North Frontal Fault System of the San Bernardino Mountains, Southern California: Cajon Pass to Lucerne Valley (Structure, Slip-Rate, Hazard), Kristian Erik Meisling, 1984 (GAX84-02492).

Photochemical Modeling of the Earth's Stratosphere, Lucien Froidevaux, Calif. Inst. of Tech., 1984 (GAX84-02489).

Prediction of Atmospheric Flow and Dispersion Over Sloping Terrain, Eueng-Nan Yeh, Univ. of Utah, 1983 (GAX84-02162).

Reaction Mechanisms and Chromatographic Behavior of Polycyclic Acid Anions in Multicomponent Ion Exchange (Arsenic, Groundwater), Liou-Liang Horng, Univ. of Houston, 1983 (GAX83-28586).

Reactions of Mutagenic Propylene Oxides with Deoxyribonucleosides and DNA, Zora Djuric, Univ. of Mich., 1983 (GAX84-02272).

Rock Geochemical Exploration at Mount Morgan, Queensland, Australia, Mark Alberti Fedikow, Univ. of New South Wales (Australia), 1982.

Sedimentary Response to Recent Tectonic Rotation in Western Oregon (Washington, Pacific

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**Northwest, Petrology.** Paul L. Heller, Univ. of Ariz., 1983 (GAX84-01264).

**Sedimentology and Invertebrate Paleontology of Triassic and Jurassic Lacustrine Deposits, Culpeper Basin, Northern Virginia.** Pamela J. Wheeler Gore, 1983 (GAX84-01325).

**Stratigraphic, Geochronological, and Petrologic Studies of the Annonioan Volcanics, North-Central Massachusetts and Southwestern New Hampshire.** John Charles Schumacher, Univ. of Mass., 1983 (GAX84-01103).

**Structure of the Brevid Zone and Blue Ridge Near Lenox, North Carolina, With Observations on Oblique Convergence and a Preliminary Theory for Irregular Structures in Shear Zones.** Anly R. Bolyarchick, State Univ. of N.Y., at Albany, 1983 (GAX84-01420).

**The Upper Proterozoic Redstone Copper Belt, Mackenzie Mountains, N.W.T., Charles W. Jefferson, Univ. of Western Ontario (Canada), 1983.**

**The Vertical Redirection of a Pollutants Tracer Due to Canine Convection.** John Andrew Rutter, Univ. of Mich., 1983 (GAX84-02364).

**Waterbed Acidification Model and the Soil Acid Neutralization Capacity Concept.** William G. Benay, McMaster Univ., Canada, 1983.

**An X-Ray Scattering and Raman Spectroscopy Study of Iron (3+), Gallium (3+) and Germanium (4+) Substituted Ammonium Sulfate Glasses.** Grant S. Heyderman, Univ. of Western Ontario (Canada), 1983.

## Geophysical Events

This is a summary of SEAN Bulletin, 9(4), April 30, 1984, a publication of the Smithsonian Institution's Scientific Event Alert Network. The complete bulletin is available in the microfiche edition of Eos as a microfilm supplement or as a paper reprint. For the microfiche, order document E84-005 at \$2.50 U.S. from AGU Fulfillment, 2000 Florida Avenue, N.W., Washington, DC 20009. For the paper reprint, order SEAN Bulletin (giving volume and issue numbers and issue date) through AGU Separates at the above address; the price is \$3.50 for one copy of each issue number for those who do not already have a deposit account. \$2 for those who do. Additional copies of each issue number are \$1. Subscriptions to SEAN Bulletin are available from AGU Fulfillment at the above address; the price is \$18 for 12 monthly issues mailed to a U.S. address, \$28 if mailed elsewhere, and must be prepaid.

## Volcanic Events

**Rabaul (New Britain):** Caldera earthquakes up 80%, two seismic crises; expansion and uplift double.

**Manam (Bismarck Sea):** Strong Strombolian activity; debris avalanches.

**Langila (New Britain):** Occasional Strombolian explosions for 10 days.

**Campi Flegrei (Italy):** Seismic energy release and uplift slow after April 1 earthquake swarm.

**Etna (Italy):** Strombolian activity and small lava flows from SE crater.

**Huone Reef (Tonga Is.):** Large pumice rafts; new island shown.

**Submarine Volcano (Izu Is.):** Acoustic waves recorded in French Polynesia.

**Macdonald (S-central Pacific):** Renewed submarine activity in 1983.

**Teahitia (French Polynesia):** Seismic swarms indicate two submarine eruptions.

**Kilauea (Hawaii):** 18th phase; four flows, longest flow of 1983-1984 eruption.

**Manna Loa (Hawaii):** Major NE Rift Zone eruption ends; total eruption volume.

**Mt. St. Helens (Washington):** Mud flow and vertical plume.

**Vesuvius (Alaska):** Vapor clouds; ash plume to 2 km altitude; no glow.

**Pagan (Mariana Is.):** Dark eruption columns. Atmospheric effects: Stratospheric aerosols decrease.

**Rabaul Caldera, New Britain Island, Papua New Guinea (4.27°S, 152.20°E).** All times are local (= UT + 8 hours).

The following is from Peter Lowenstein. "A further intensification of seismic activity in the Rabaul Caldera took place in April. The total number of caldera earthquakes was 13,749, 60% more than in March (the March total was 8729; see last month's Bulletin). Seismicity was concentrated on the E side of the caldera, in Greet Harbour and at the entrance to Blanche Bay."

"Major seismic crises occurred on April 21 and 22, when 1011 and 1717 events were recorded. The crisis on April 21 was centered at the mouth of Blanche Bay, and the strongest earthquake was a magnitude (ML) 3.8 event. Only minor ground deformation was associated with this crisis."

"On April 22 at 1100 an ML 4.8 earthquake heralded the most energetic crisis in date, which was centered at the head of Greet Harbour. Structural damage in this and the Sulphur Creek area included cracking, and in one case, collapsing, of masonry walls, cracks in concrete floors, a burst water main, and burst household water tanks. This around Greet Harbour ranged from 90 to 80 micro-radial inflation centered in the Harbour. Measurements of horizontal deformation indicated expansion of the Greet Harbour area by 20-30 microradians."

"The overall pattern of ground deformation in April indicated that the strongest tilting, of up to 80 microradians, was in the Greet Harbour area. Rates of horizontal deformation indicated expansion was about

double that in any previous month (40-50 microradians).

"Leveling surveys from Rabaul Township to Matupit Island and around Greet Harbour showed that between mid-March and mid-April the S end of Matupit Island rose 76 cm. Further uplifts of about 50 cm on Matupit, and at the head of Greet Harbour, accompanied the April 22 seismic crisis, making the total uplift in April about double that in any previous month."

Information Contact: Peter Lowenstein, Principal Government Volcanologist, Rabaul Volcano Observatory, P.O. Box 386, Rabaul, Papua New Guinea.

**Home Reef Volcano, Tonga Islands, S Pacific (18.59°S, 174.78°W).** All times are local (= UT + 13 hours).

An early March eruption of Home Reef produced large quantities of pumice, ejected an eruption cloud to more than 12 km altitude, and built a new island (see Eos, March 1983).

Tonga government geologist David Tappin reported that brown discolored water preceded the eruption, which started March 1-2. The new island was visible by March 2. When Captain Jeff Heard of South Pacific Islands Airways flight 807 over the eruption site on March 5 at 1030, explosive activity had declined. Weak steaming occurred from a submarine crater surrounded by the new island.

In mid-March, a cargo vessel traveling from Tonga to Fiji at 12 km per hour took 9 hours to pass through a zone of pumice. Samples were collected from this vessel about 10 km W of Tonga. Pumice rafts were reportedly sighted at Oneata Island, Lau Group (18.45°S, 178.50°W, roughly 500 km WNW of Home Reef) on April 5. On May 1, ships between Tonga, Fiji, and Samoa reported that floating pumice was so thick that it was clogging their seawater intake systems.

Personnel from the Royal New Zealand Air Force (RNZAF) flew over the new island March 23. They gave its location as 19.02°S, 174.73°W, about 10 km S of Late Island. Dimensions of the new island were estimated at 1500 m by 500 m, with cliffs about 50-60 m

high. Discolored water just NW of the island suggested submarine activity. Photographs taken from upwind showed the island to be yellowish brown in color, but atmospheric haze caused it to appear dark brown from downwind. David Tappin reported that activity was continuing in early April.

The Réseau Sismique Polyésien (RSP) did not record any seismicity from the eruption. Islands and deep water between Tahiti and Tonga prevented RSP stations from recording any acoustic waves (T-phase).

Information Contacts: David Tappin, Government Geologist, Nukunono, Tonga; Warant Officer P. J. R. Shepherd, 1 SQN ALM LDR, RNZAF Whenuapai, Auckland, New Zealand; J. H. Lattier, DSIR Geophysics Division, P.O. Box 1320, Wellington, New Zealand; J. Lum, Ministry of Energy and Mineral Resources, Private Mail Bag, Suva, Fiji; Ram Krishna, Director of Meteorology, Fiji Meteorological Service, Private Mail Bag, Nandi Airport, Fiji; J.M. Talandier, Directeur, Laboratoire de Géophysique, Commissariat à l'Energie Atomique, B.P. 640, Paapeete, Tahiti, Polynésie Française.

Information Contact: National Earthquake Information Service, U.S. Geological Survey, Stop 967, Denver Federal Center, Box 25066, Denver, CO 80225 USA.

**Meteoritic Events**

Fireballs: Papua New Guinea; Hawaii, Mississippi River Valley, USA.

**Correction**

The names of the people who wrote the tribute for Mahdi S. Hantush, which appeared May 29, 1984, were inadvertently left off of the article. They are M. U. Ahmad, Ohio University, Athens, Ohio; G. W. Gross, New Mexico Institute of Mining and Technology, Socorro, New Mexico; M. A. Marino, University of California, Davis, California; S. S. Papadopoulos, S. S. Papadopoulos and Associates, Inc., Rockville, Maryland; and Z. A. Saleem, Ebasco Services, Inc., Greensboro, North Carolina.

**Earthquakes**

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high-frequency spasmodic and harmonic tremor. Teahitia, a seamount with a summit about 2 km below sea level, was the site of seismicity associated with submarine eruptions detected by the RSP in March-April 1982 and July 1983 (see SEAN Bulletin, 7(4), and 8(8)).

Information Contact: J.M. Talandier, Directeur, Laboratoire de Géophysique, Commissariat à l'Energie Atomique, B.P. 640, Paapeete, Tahiti, Polynésie Française.

**Earthquakes**

Information Contact: National Earthquake Information Service, U.S. Geological Survey, Stop 967, Denver Federal Center, Box 25066, Denver, CO 80225 USA.

**Meteoritic Events**

Fireballs: Papua New Guinea; Hawaii, Mississippi River Valley, USA.

**Correction**

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**Postdoctoral Position/UCLA.** Postdoctoral position in experimental geophysics/petrology available immediately for research on upper-mantle or lower-crustal problems. Successful applicant will have a strong background in thermodynamics and petrology. Send application to: Dr. Robert E. Anderson, Institute of Geophysics and Planetary Physics, University of California, Los Angeles, California 90024, telephone (213) 825-8868.

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**Assistant Research Geophysicist.** The Institute of Geophysics and Planetary Physics or the Ocean Research Division of the Scripps Institution of Oceanography are considering the appointment of an assistant research geophysicist, at the level of a research group conducting electromagnetic soundings of the ocean floor. Applicants should have experience with land and ocean EM measurements, a demonstrated capacity to design and construct equipment, and the ability to carry out experiments at sea. A Ph.D. in geophysics or related sciences is required. Candidates should have some experience with the analysis and interpretation of EM data. Salary range is \$25,100-\$35,100. Applicants must submit a resume, copies of relevant publications, and the names of three references by 1 July 1984 to:

Dr. Alan Chave  
University of California, San Diego  
Institute of Geophysics and Planetary Physics  
La Jolla, CA 92093.

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**Microgeology/University of Puerto Rico, Mayaguez.** Position open July 1, 1984. Assistant Professor level, tenure track, salary \$17,800 per annum (9 months teaching). Ph.D. required. Duties will involve teaching at the graduate level courses in the discipline being considered here, supervising student research and conducting personal research. Applicants should send curriculum vitae, a brief statement of teaching and research plans and three letters of recommendation to: Chairman, Appointment Committee, Department of Marine Sciences, University of Puerto Rico, Mayaguez, P.R. 07080. Telephone 809-832-0400, ext. 5143.

**Research Position-Space Physics/Rice University.**

The Space Physics and Astronomy Department at Rice University seeks applicants for one or more full-time research positions within the department. Successful applicants will play key roles in the development of theoretical three-dimensional models of the Earth's electromagnetic field. Applicants should have knowledge of, and interest in, at least one of the following areas: solar-wind magnetosphere interactions, magnetosphere-ionosphere coupling, ionosphere-atmosphere coupling, collisionless plasma microphysics, atmospheric electricity. Experience and/or interest in numerical modeling is an important consideration.

Rate and salary level commensurate with experience, ranging from one-year Research Associateship (renewable in subsequent years depending on performance) to open-ended Research Scientist appointment in the Center for Space Physics. Please send resume and names of three professional references to: T. W. Hill or R. A. Wolf, Space Physics and Astronomy Dept., Rice University, Houston, TX 77251.

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**University of Cambridge/Theoretical Seismologist.** It is hoped soon to appoint a postdoctorate to work independently in the general field of theoretical seismology. An interest in seismic modeling and interpretation, particularly of body-waves, would be suitable. Stimulating environment with other theoretical, reflection, reflection and earthquake seismologists. University salary. Send curriculum vitae to Professor C.H. Chapman, Bullard Laboratories, Department of Earth Sciences, University of Cambridge, Madingley Road, Cambridge CB3 0ET, England, by 31 July 1984.

**Research Associate/Research Technician.** The University of Maine at Orono (UMO) has an opening for a research associate/research technician who would work in a small geophysical group. We seek an individual who can use and maintain modern digital electronic equipment; for example, multi-channel analyzers, VU interfaces for microcomputers, digital plotters and digitizing tablets. Familiarity with BASIC and FORTRAN will be needed, and some geophysical field work may be required as part of the duties of the appointee. Current funding permits an appointment for at least 12 months. Subject to arrival of anticipated funding, the appointment period could be extended to two years, or longer. Call Edward R. Decker at 207-581-2158 or 207-581-2152 about the position. Otherwise, send inquiries, a resume and a list of at least three references to Edward R. Decker, Department of Geological Sciences, 110 Boardman Hall, University of Maine at Orono, Orono, ME 04469.

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**Research Associate/Brown University.** Research Associate in Planetary Geology at Brown University, Providence, Rhode Island. Experience in geologic/geomorphic analysis of planetary images, study of surface geologic processes, computerized image processing, and/or quantitative geomorphology is desirable. Deadline for applications is June 30, 1984. Submit resume, names and addresses of three references to Dr. James H. Neal, Box 1846, Brown University, Providence, RI 02912.

Brown University is an equal opportunity/affirmative action employer.

**Postdoctoral Research Associate Positions/Geophysics and Igneous Geochemistry.** The University of Maine at Orono (UMO) has postdoctoral openings for a solid earth geophysicist and an igneous geochemist. We seek a geophysicist who wishes to advance fundamental understanding of past and current thermal histories of the Appalachian Orogen in New England and elsewhere. The geochemist would be expected to investigate volcanic and plutonic rocks in the Appalachian Orogen in Maine and in other terranes. Current funding permits appointments for at least 12 months. Subject to arrival of anticipated funding, the appointments could be extended to two years. Send resumes and names of three references to Dr. James H. Neal, Box 1846, Brown University, Providence, RI 02912.

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**Assistant Professor of Geology.** A one term, renewable (possible tenure track) position begins September 1, 1984. Ph.D. is required; needs strong background, interest in sedimentary geology. Will teach graduate-level courses in sedimentary geology and have strong interest in research. Salary commensurate with qualifications and experience. Review of credentials begins June 30, 1984. Send via, statement of teaching and research interests and names of four references to: Search Committee, Department of Geology, Eastern Washington University, Cheney, WA 99004.

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**Research Associate/University of Maryland.** The Space Physics Group of the Department of Physics and Astronomy has an opening for a Research Associate beginning as early as July 1, 1984 for an initial one-year period with high likelihood of extension. The position involves research on energetic particles of solar and interplanetary origin. Applicants should possess a Ph.D. in a relevant area of physics or astrophysics relevant research experience is highly desirable. Inquiries and applications should be addressed to Professor Glenn M. Mason, Department of Physics and Astronomy, University of Maryland, College Park, MD 20742. Applicants should send a resume, a statement of research interests and a description of research experience, and should arrange for the sending of at least three letters of reference.

The University of Maryland is an equal opportunity/affirmative action employer.

**Postdoctoral Fellow in Igneous Petrology.** Available August 15, 1984, duration of 1-2 years. Areas of research include mineralogy/petrology/geochemistry of igneous rocks. A working knowledge of the electron microprobe is required. Please send resume, short summary of research goals and the names of three persons who may be contacted for recommendation to: L.A. Taylor, University of Tennessee, Department of Geological Sciences, Knoxville, TN 37996. Telephone: 615-974-6013.

**Postdoctoral Research Positions in Planetary Atmospheres/Lunar and Planetary Laboratory, University of Arizona.**

Two positions will involve research in planetary physics and analysis of UV data from the Voyager mission. Research opportunities for these positions include the bound and extended atmospheres and ionospheres of the giant planets and their satellites, the ionosphere of Venus, and the atmosphere of Mars. Applicants should have a strong background in theory and data analysis. Physicists and astronomers are encouraged to apply. Curriculum vitae, bibliography and three letters of reference should be sent by July 15, 1984, to: Dr. A. L. Broadfoot, Lunar and Planetary Laboratory, University of Arizona, 3625 E. Ajo Way, Tucson, Arizona 85713.

The University of Arizona is an Equal Opportunity Employer.

**Electronic Engineer.** Technical and supervisory position in new electronics facility in University of Arizona's Lunar & Planetary Laboratory. Design, construction, and maintenance of specialized scientific instruments and computer-based systems for laboratory, astronomical, and spacecraft applications. Minimum requirement of Bachelor's Degree and 5 years experience or equivalent. Experience with military or space-related hardware desirable. Send resume to: Staff Employment Office, University of Arizona, 1717 E. Speedway Blvd., Tucson, AZ 85721.

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#### STUDENT OPPORTUNITIES

**Special Doctoral Research Assistantships.** The Department of Oceanography of Old Dominion University has several special doctoral research assistantships available for Fall Semester, 1984 and 1985. These carry a stipend of \$7,000 per academic year, renewable for three years. Applicants with M.S. degrees qualify for waiver of tuition. Students interested in obtaining the Ph.D. in the area of biological, chemical, geological, or physical oceanography should send an introductory resume to Dr. Ronald E. Johnson, Graduate Program Director, Department of Oceanography, Old Dominion University, Norfolk, VA 23508.

Old Dominion University is an affirmative action/equal opportunity institution.

## Meetings

### Announcements

#### Water Quality Modeling

September 6-7, 1984 Conference on Stormwater and Water Quality Management Modeling, Hamilton, Ontario, Canada. (William James, Civil Engineering Department, McMaster University, Hamilton, Ontario, Canada, L8S 4L7.)

The deadline for submitting abstracts is July 1, 1984.

Conference talks will concern modeling pollution from urban and industrial sources. Papers are particularly solicited on emerging microcomputer hardware and software techniques in stormwater, flood, and pollution management.

#### Tar Sand Symposium

June 26-29, 1984 WRI-DOE Tar Sand Symposium, Vail, Colo. Sponsors: Western Research Institute, U.S. Department of Energy, (L. C. Marchant, Western Research Institute, P.O. Box 3305, University Station, Larimer, WY 82071; tel: 307-721-2903.)

Technical sessions will run from June 27-29, with informal technical discussions planned for June 28. Among the topics addressed by formal speakers will be resources, in situ extraction processes, bitumen properties and upgrading, mining and plant extraction processes, and environmental and economic concerns.

#### Environmental Data Users

August 19-22, 1984 Pathways and Future Directions for Environmental Data and Information Users, Denver, Colo. Sponsor: NOAA, (SES, Incorporated, P.O. Box 2697, Springfield, VA 22152.)

Early registration deadline is July 1, 1984.

The objective of the conference is to increase user awareness of the range of environmental data available to the public. The program will consist of panel presentations followed by question and answer periods. The use of environmental data in many disciplines will be addressed, including legal and insurance industries, travel and recreation industries, communications and aerospace industries, urban and regional planning, renewable resources management, health and natural hazard monitoring, architecture and engineering, and marine exploration and coastal zone development. There will also be exhibits and displays from each of the four data centers of the National Environmental Satellite, Data, and Information Service.

#### Magnetism and Magnetic Materials

November 27-30, 1984 Thirtieth Annual Conference on Magnetism and Magnetic Materials, San Diego, Calif. Sponsors: American Institute of Physics, Magnetics Society of IEEE, (John Scott, American Institute of Physics, 355 East 45th St., New York, NY 10017.)

The deadline for submitting abstracts is July 6, 1984.

The purpose of the conference is to bring together scientists and engineers interested in recent developments in all branches of fundamental and applied magnetism. Emphasis is placed on experimental and theoretical research in magnetism, the properties and synthesis of new magnetic materials, and advances in magnetic technology. The program will consist of both invited and contributed papers.

### Meeting Report

#### Yosemite Conference

A conference on Planetary Plasma Environments: A Comparative View was held in Yosemite, Calif., January 30 to February 3, 1984. The purpose of the conference was to discuss the comparative aspects of planetary and cometary plasma systems. Major support for the conference was provided by the National Aeronautics and Space Administration, Office of Space Science and Applications. Additional sponsorship was provided by Ball Aerospace Systems Division, TRW Space and Technology Group, the American Geophysical Union, and Stanford University.

Over the last 10 years, we have obtained important new information on the planetary plasma environments of Venus, Jupiter, and Saturn. In addition, there are several planned cometary flyby and rendezvous missions which will be taking a first in situ look at these interesting objects in the next 5 years. This,

coupled with the increased exploration and understanding of the earth's system (particularly the study of magnetosphere-ionosphere coupling, which has been the subject of several past Yosemite meetings), made the topic of comparative planetary plasma environments an excellent choice for the fifth biennial Yosemite topical conference.

The program was organized into five main sessions. The opening session provided an overview of the major morphological features of planetary and cometary systems. The following two sessions discussed the plasma and energy sources that operate in these systems. These early sessions provided the framework for the final two sessions, which dealt with the interaction of the various plasma components of the systems and with the commonality of the physical processes that take place within the solar system environments.

Chris Russell of the University of California, Los Angeles began the first session with a whirlwind tour discussing the solar wind interactions with various inner solar system bodies. The outer planets were then addressed by Ralph McNutt of the Massachusetts Institute of Technology, and Tom Cravens of the University of Michigan discussed the interaction of comets with the solar wind.

The first session identified several interesting questions that were explored in subsequent sessions of the conference: (1) What is the nature of the solar wind interaction with Mars? (2) How important is the role of ionospheric plasma as a source of magnetospheric plasma at the outer planets? (3) What is the difference between the solar wind/magnetosphere and solar wind/Venus interactions? These questions not only stimulated much discussion throughout the conference, but also suggested future planetary missions and instrumentation that could help address these questions.

The second session focused on sources of plasma within planetary and cometary systems. Doug Torr discussed the role of both electron precipitation and solar EUV production of ionospheric plasma with emphasis on the need for proper calibration of the existing solar EUV data sets. He indicated that in regions of the EUV spectrum there may be as much as a factor of 3 error in the widely used Hinteregger EUV fluxes and that proper calibration can help to explain the long standing discrepancy between the modeled and measured photoelectron spectrum.

Tamas Gombosi contrasted and compared the structure of the Venusian, Martian, and terrestrial ionospheres. While O<sup>+</sup> is the major ion in the upper ionosphere for all three planets, the minor ion composition is weighted toward nitrogen-derived ions for Venus and carbon dioxide-derived ions for Mars. It was also noted that the maintenance of the nighttime Venusian ionosphere is due mainly to ion flows from the dayside driven by plasma pressure gradients, rather than particle precipitation as first suspected from early Soviet Venera data.

The ionospheric structure in the outer planets was reviewed by Sushil Atreya. He indicated that in the high latitude ionospheres of Jupiter, Saturn, and Uranus the auroral precipitation energy is large enough to provide the dominant source of ionization but that at low and mid-latitudes, solar EUV sources were thought to dominate.

Asoka Mendis identified four possible sources of ions for cometary systems: (1) photoionization, (2) charge exchange with solar wind ions, (3) electron impact ionization, and (4) critical velocity ionization. He reviewed both chemical and dynamical modeling of comets, noting some of the controversial ideas that have developed from these models. These ideas fueled the discussion that followed the talk. The discussion centered on the role of friction, between the dust and the escaping gas in a cometary system, as a means of providing a choke for a transonic gas expansion in the cometary atmosphere.

The solar wind as a source of ionization in magnetospheric systems was discussed by Tim Eastman. Tim estimated that about 7 x 10<sup>10</sup> ions s<sup>-1</sup> enter the solar wind source sphere from the solar wind source that in quiet times the solar wind source dominated the plasma sheet composition, yet in storm times the earth's ionosphere became the dominant source of plasma sheet plasma. He suggested that effusive processes associated with wave particle interactions is a major process for solar wind plasma entry to the earth's magnetosphere and that magnetic merging is less significant. Tom Hill disagreed with this point of view during the discussion period, suggesting that the relative importance of the two processes is still an open question.

Critical velocity ionization was discussed by Patrick Newell of the University of California, San Diego. Although Titan, Io, and comets have been suggested as locations where this effect might be an important source of ionization, very little experimental evidence for critical velocity ionization in space plasmas exists at this time. Peter Banks, nevertheless, suggested that this mechanism might be responsible for the high ion densities (6 x 10<sup>10</sup> cm<sup>-3</sup>) seen around the space shuttle during

dayside orbital segments on STS-5.

The energy sources session next focused attention on the energization of plasmas in planetary systems. Bill Knudsen began the session by discussing plasma energization at the inner planets with particular emphasis on Venus using Pioneer Venus measurements. He indicated that as one crosses from the solar wind through the mantle interaction region to the Venusian ionosphere, the electron temperature goes from 10 to 100 eV due to ion exchange with the solar wind. Waves are also produced by the solar wind interaction that, through Landau damping, serve as a heat source for the topside ionosphere. Knudsen estimated that about 5% of the solar wind energy incident on Venus comes into the ionosphere.

This behavior of a fairly direct energization of the planetary plasma at Venus was contrasted with more complex solar wind/magnetosphere energization processes in a review by Michael Schulz. The observed monotonic decrease of energetic phase space density with decreasing  $L$  indicates that the plasma energization in the earth's magnetosphere is intrinsic to the inward transport driven by solar wind/magnetospheric convection.

Alex Dessler gave an interesting talk, suggesting that planetary rotation, by exerting a magnetic torque, may be a general source of energy for magnetosphere systems. He indicated that there are two ways to supply the torque: (1) to have an internal plasma source such as Io in the Jovian system or (2) to have an external plasma source such as the solar wind. Alex and coworkers have done some recent calculations, indicating that this may be an important source of energy for the magnetospheres of Jupiter, Saturn, Uranus, and possibly Neptune.

In the absence of collisions within planetary plasma systems, waves provide the required dissipation and thus serve as important means of redistributing energy within the system. Maha Ashour-Abdalla suggested that there are two types of wave heating: (1) quasi-linear heating by many low amplitude wave modes, or (2) heating by a few strong wave modes where particle orbits that were originally bounded become unbounded and stochastic. The remainder of her presentation dealt with a specific example of the latter process.

While the first three sessions dealt with the separate elements of planetary and cometary plasma systems, the fourth session (interactive plasma processes) addressed the interplay between the separate components. The session began with a presentation of some examples of terrestrial magnetosphere-ionosphere coupling processes.

Hunter Waite reviewed recent results from the Dynamics Explorer satellite program. These included recent studies on ion-neutral coupling processes and also DE 1 observations of low-energy ion outflows observed at high altitudes in the terrestrial magnetosphere. In particular, significant fluxes of O<sup>+</sup> are commonly observed flowing out of the ionosphere at high altitudes over the polar cap.

A review of terrestrial ionosphere-magnetosphere modeling was given by Dick Wolf. Carl McIlwain questioned the substorm scenario and suggested that the magnetic field configuration might simply set up the electric field that then led to auroral precipitation. Carl McIlwain and Peter Banks both asked if it was possible that mass loading of the high latitude magnetosphere was an important consideration for the convection electric field, a question certainly inspired from earlier presentations on mass loading of the solar wind at Venus and in comet systems.

Janet Luhmann presented results on the Venus-solar wind interaction. She indicated that there is a conversion of solar wind dynamic pressure to magnetic pressure that is matched by the ionospheric pressure at the ionopause. The solar wind dynamic pressure controls the altitude of the ionopause. When the altitude approaches its minimum observed value of 200 km, large-scale magnetic fields develop within the ionosphere. These fields, according to Luhmann, can be explained as a result of a combination of vertical convection of magnetic flux from the ionopause and diffusion. Rob Wolfe, however, asserted that the Cloutier model of the interaction does not agree with this interpretation. In the Cloutier model, a large-scale current system is driven in the ionosphere by absorption of solar wind plasma. A long discussion on the virtues of both model ensued, and it was clear that more careful study will be required to determine which physical scenario produces the correct magnetic field within the Venus ionosphere.

The emphasis then switched to the outer planets and the interesting new results that show the importance of moons and rings as sources and sinks of magnetospheric plasma. Chris Goertz categorized the effects of moons by whether they had an atmosphere and whether they were magnetized or not.

The interaction of comets with the solar wind was revisited by W. H. Ip. Again, the contrast of the cometary case with the Venus interaction focused on the greater degree of mass loading ahead of the shock and possible strong wave particle interactions that could lead to the formation of a weak cometary bow shock. Ip also suggested that the wavy or kink configurations observed in cometary ion tails are indicative of Kelvin-Helmholtz or

Rayleigh-Taylor instabilities triggered by solar wind disturbances. In addition, Rick Elphic presented some evidence of filamentary structures in the Venus nightside ionosphere that are quite suggestive of comet ion tails and may serve as further evidence of similar physical processes in the Venus and cometary-solar wind interactions.

The final session attempted to identify and describe some of the common plasma processes that occur in solar system plasmas. Ron Lepping began the session with a talk on how shock and magnetopause formation. All planets observed to date have a bow shock and all but Venus have a magnetopause. The characteristics of the bow shocks of planets depend strongly on the solar wind conditions, whereas the shape of a magnetopause will depend crucially on the detailed three-dimensional pressure profile it presents to the solar wind.

Acceleration processes in plasma systems appear to be an important common process as demonstrated by the observation of auroras at Earth, Jupiter, Saturn, and Uranus.

Tom Hill opened his talk on the subject by indicating that acceleration of particles in a magnetosphere implied the violation of the rules of MHD. He classified these processes according to the invariant violated and according to the timescales or their physical mechanism. He concluded by pointing out three new challenging questions for future contemplation: (1) What produces energetic particle bursts? (2) Why does the Io plasma torus heat as it moves radially outward from Jupiter? (3) Why is O<sup>+</sup> accelerated in the terrestrial polar cap?

Transport processes within plasma systems were reviewed by George Siscoe. Charged particles move through magnetospheres in three ways: (1) convection, (2) diffusion, and (3) field-aligned winds. Charge exchange, however, also causes mass transport by converting magnetically bound ions into fast unbound neutrals. Steady and nonsteady convection is driven in the earth's magnetosphere by merging between the planetary magnetic field and the solar wind magnetic field. This is to be contrasted with the magnetosphere of Jupiter, where convection may be driven by centrifugal force acting on an inhomogeneous distribution of corotating plasma. Siscoe also noted that when induced electric fields are present, as must be the case during magnetic field collapse or inflation, magnetospheric motions are not necessarily reflected in ionospheric motions. Nonsteady magnetic merging can also generate plasma bubbles (plasmoids) which are a source of field-aligned winds. Diffusion is needed to account for particles with energies higher than the convection potential and is important in the earth's radiation belts and possibly in Jupiter's middle magnetosphere. Transport processes can also be driven by atmospheric circulation. Indeed, in a magnetosphere, stresses on both ends of a magnetic field line must be balanced and communicated by means of field-aligned currents.

Meetings (cont. on p. 396)

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